

The initial value problem for a massless scalar field and N point-charges in one space dimension

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What is a Classical Point Charge?

- On $n + 1$ dimensional space-time $\mathbb{R}^{1,n} \cong \mathbb{R} \times \mathbb{R}^n$ with metric $\eta = \text{diag}(1, -1, -1\dots)$ and coordinates $x = (x^0, x^1, x^2, \dots)$.

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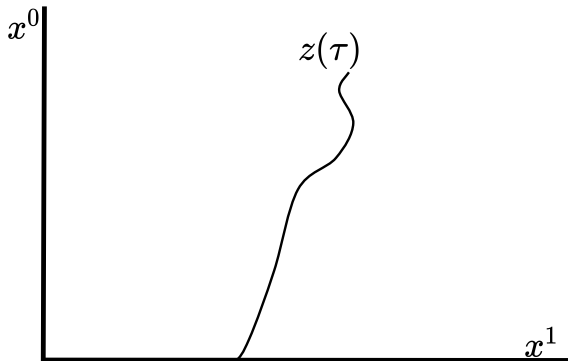
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- Notation: Jump and Average:

$$\bar{f}(z) = \lim_{\epsilon \rightarrow 0} \frac{f(z + \epsilon) + f(z - \epsilon)}{2}, \quad [f]_{x=z} = \lim_{\epsilon \rightarrow 0^+} (f(z + \epsilon) - f(z - \epsilon)) \quad (1)$$

Scalar Point Charge

The scalar point charge acts as a source for the scalar field via

$$\square U := \partial_0^2 U - \Delta U = a \int \delta^{(n+1)}(x - z(\tau)) d\tau \quad (2)$$



Motion of Scalar Field Animation

Given initial data (V_0, V_1) and a charge trajectory $z(\tau)$, the solution for U is well-known.

$$\begin{cases} \square U = a \int \delta^{(n+1)}(x - z(\tau)) d\tau \\ U|_{x^0=0} = V_0 \\ \partial_0 U|_{x^0=0} = V_1 \end{cases} \quad (3)$$

Figure: Animation of scalar field sourced by a moving point charge in one space dimension.

(In)admissible Force Laws

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- In one space dimension, no modification is necessary!

Conservation of Stress Energy-Momentum

Theorem

Suppose U satisfies (3) on $\mathbb{R}^{1,1}$ with “admissible” initial data $(U|_{x^0=0}, \partial_0 U|_{x^0=0})$ and $z(\tau)$ is an arbitrary time-like world-line.

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This force can be explicitly computed as

$$\frac{dp^\nu}{d\tau} = -a \overline{\partial^\nu U}. \quad (11)$$

Derivation of Force Law 1

$$\int_{\Omega} \partial_{\mu} T^{\mu\nu} dV = \int_{\partial\Omega} T^{\mu\nu} N_{\mu} dS \quad (12)$$

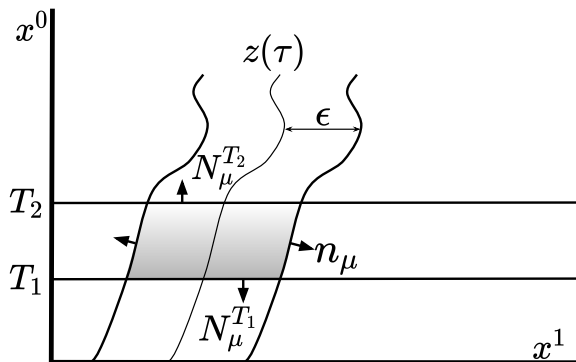
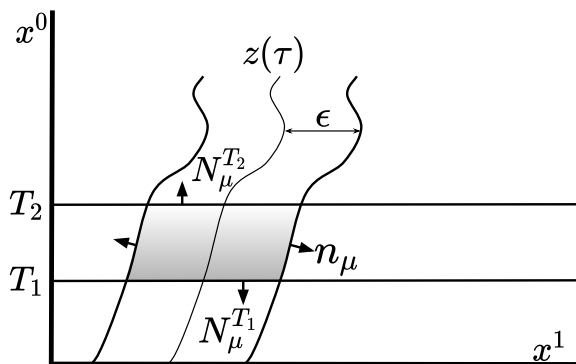


Figure: Region of integration Ω and its normals

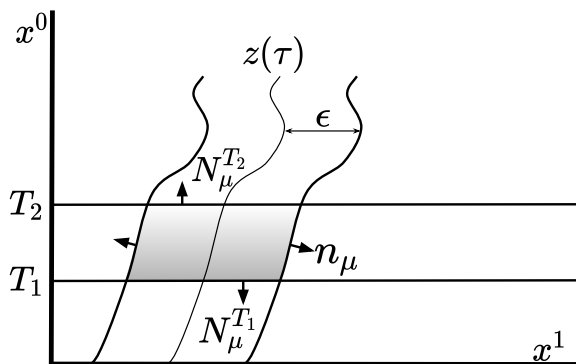
Derivation of Force Law 1

$$\int_{\partial\Omega} T_p^{\mu\nu} N_\mu dS = p^\nu(\tau_2) - p^\nu(\tau_1) = \int_{\tau_1}^{\tau_2} \frac{dp^\nu}{d\tau} d\tau. \quad (12)$$



Derivation of Force Law 1

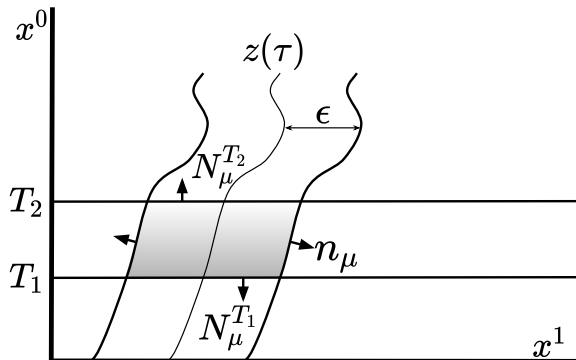
$$\int_{\tau_1}^{\tau_2} \frac{dp^\nu}{d\tau} d\tau = - \int_{\partial\Omega} T_U^{\mu\nu} N_\mu dS. \quad (12)$$



Derivation of Force Law 2

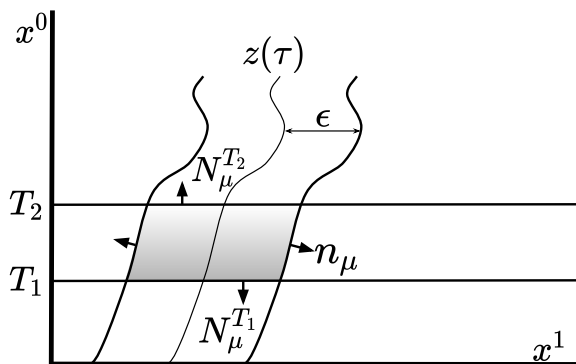
Lemma

Suppose U satisfies equation (3) with $(U, \partial_0 U)|_{x^0=0} \in C^{0,1}(\mathbb{R}) \times L^\infty(\mathbb{R})$ and singular source concentrated on a time-like world-line $z(\tau)$. Then U is of regularity $C^{0,1}(\mathbb{R}^{1,1})$, and $\partial_\mu U, T_U^{\mu\nu} \in L^\infty(\mathbb{R}^{1,1})$.



Derivation of Force Law 2

$$\lim_{\epsilon \rightarrow 0} \int_{\partial\Omega} T_U^{\mu\nu} N_\mu dS = \int_{\tau_1}^{\tau_2} [T_U^{\mu\nu} n_\mu(z^0, x^1)]_{x^1=z^1(\tau)} d\tau \quad (13)$$



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- The final force law $\frac{dp^\nu}{d\tau} = -a\overline{\partial^\nu U}$ follows from

$$\left[\frac{f^2(x^1)}{2} \right]_{x^1=z^1} = [f]_{x^1=z^1} \bar{f} \quad (16)$$

Force Law Consequences 1

Corollary

The force contribution of external scalar fields follows the standard law derived from the principle of least action.

$$F_{\text{ext}}^\nu = -a\partial^\nu U_{\text{ext}}. \quad (17)$$

However, the singular “self-force” can now be determined, and the expression for it guarantees the conservation of the system’s total energy-momentum.

$$F_{\text{self}}^\nu = -a\overline{\partial^\nu U_{\text{source}}} = \frac{-a^2 u^\nu}{2} \quad (18)$$

Theorem

For any set of particle parameters with positive bare mass and non-zero real scalar charge, and for any “admissible” initial data, the joint initial value problem given by

$$\begin{cases} \eta^{\mu\nu} \partial_\mu \partial_\nu U(x) &= a \int \delta^{(2)}(x - z(\tau)) d\tau \\ U(0, x^1) &= V_0(x^1) \\ \partial_0 U(0, x^1) &= V_1(x^1), \end{cases} \quad (19)$$

$$\begin{cases} \frac{dz^\nu}{d\tau} = u^\nu \\ \frac{dp^\nu}{d\tau} = - [n_\mu T_U^{\mu\nu}(z^0, x^1)]_{x^1=z^1(\tau)}. \end{cases} \quad (20)$$

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Force Law Consequences 3

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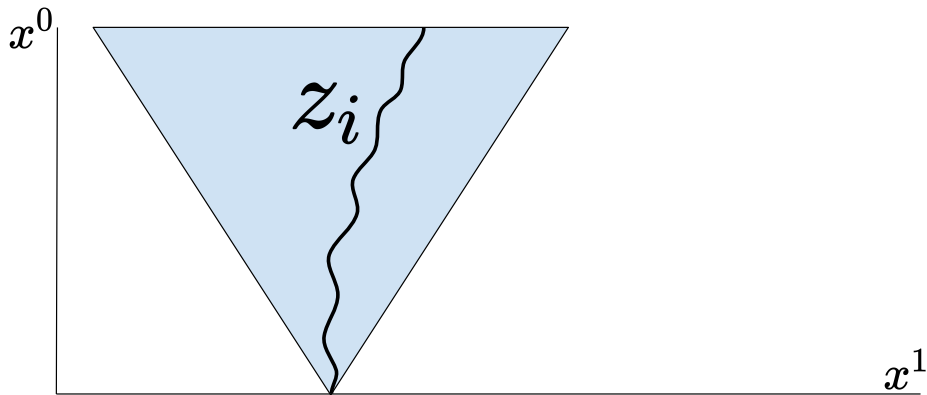
$$\begin{cases} \eta^{\mu\nu} \partial_\mu \partial_\nu U(x) = \sum_{j=1}^N a_j \int \delta^{(2)}(x - z_j(\tau)) d\tau \\ U(0, x^1) = V_0(x^1) \\ \partial_0 U(0, x^1) = V_1(x^1), \end{cases} \quad (21)$$

$$\begin{cases} \frac{dz_j^\nu}{d\tau} = u_j^\nu \\ \frac{dp_j^\nu}{d\tau} = - [n_\mu T_U^{\mu\nu}(z^0, x^1)]_{x^1=z_j^1(\tau)}. \end{cases} \quad (22)$$

admits a unique, local-in-time solution until any particle reach the speed of light or any two particles cross.

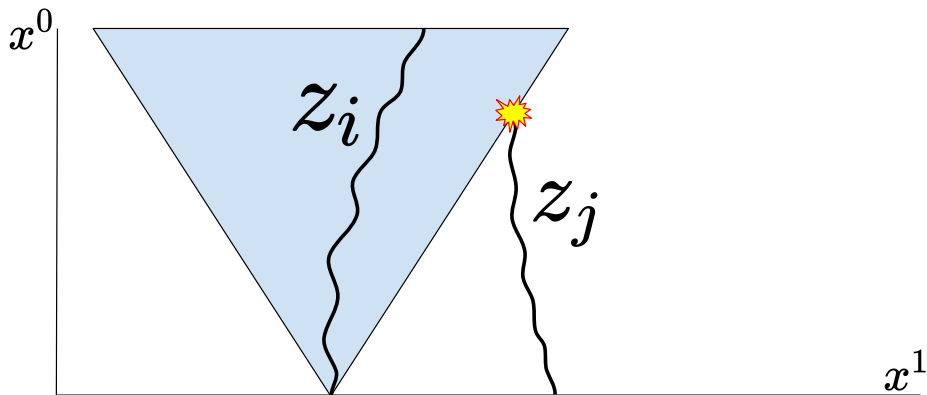
Admissible Initial Data

$$(U, \partial_0 U)|_{x^0=0} \text{ smooth} \Rightarrow \partial_\mu U \text{ jumps across } \mathcal{F}^+(0, z_i^1(0)) \quad (23)$$



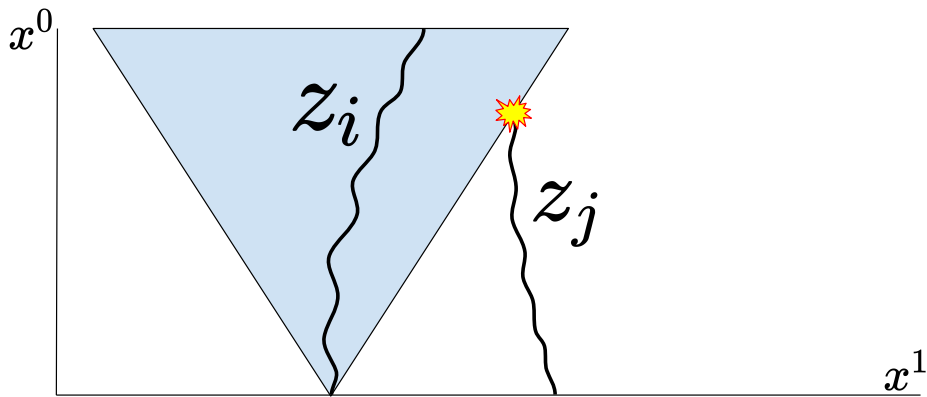
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Admissible Initial Data

$$(U, \partial_0 U)|_{x^0=0} \text{ smooth} \Rightarrow \text{lose local well-posedness fast!} \quad (23)$$



Theorem

Given scalar field initial data $(U, \partial_0 U)|_{x^0=0}$ for which $\partial_1 U|_{x^0=0}$ has N “jumps”, the joint initial value problem given by (21) and (22) admits classical solutions for $z_i(\tau)$ if and only if a point charge source is initially placed at each jump in $\partial_1 U|_{x^0=0}$, and the following equations hold at these locations

$$\begin{cases} [\partial_\mu U|_{x^0=0}]_{x^1=z_i(0)} \text{ is space-like} \\ a_i = \eta^{\mu\nu} [\partial_\mu U|_{x^0=0}]_{x^1=z_i(0)} [\partial_\nu U|_{x^0=0}]_{x^1=z_i(0)} \\ (u_i^0(0), u_i^1(0)) = \frac{1}{a_i} [(\partial_1 U, -\partial_0 U)|_{x^0=0}]_{x^1=z_i(0)} \end{cases} \quad (23)$$

1-D Gravity

Summary

- We studied the joint evolution equations for a scalar field and its point charge source in 1 space dimension

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- Admissibility condition: Initial field data uniquely determines $(a_j, z_k(0), u_j^\mu(0))!$

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- Point mass acts as singularity in curvature

$$g_{\mu\nu} \in C^{0,1}(\mathcal{M}), \quad [\Gamma_{\mu\nu}^{\kappa}(z^0, x^1)]_{x^1=z^1} \neq 0 \quad (24)$$

$$\nabla_{\mu} T_p^{\mu\nu} \propto \partial_{\mu} T_p^{\mu\nu} + \Gamma T_p = 0? \quad (25)$$

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- True force law from Energy-Momentum conservation:

$$\int_{\tau_1}^{\tau_2} \frac{dp^\nu}{d\tau} d\tau = \lim_{\epsilon \rightarrow 0^+} \left(\int_{\tau_1}^{\tau_2} n_\mu T_U^{\mu\nu}(z^0, z^1 + \epsilon) - n_\mu T_U^{\mu\nu}(z^0, z^1 - \epsilon) d\tau \right) \quad (26)$$

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- Action Principle:

$$S[U, z] = \int (\mathcal{L}_U + \mathcal{L}_i + \mathcal{L}_p) \sqrt{-\eta} dx^2 \quad (27)$$

$$\mathcal{L}_U := \frac{1}{2} \eta^{\mu\nu} \partial_\mu U \partial_\nu U, \quad \mathcal{L}_i := \frac{-a}{m} U(z) \mathcal{L}_p \quad (28)$$

$$\mathcal{L}_p := -\frac{m}{\sqrt{-\eta}} \int \sqrt{\eta_{\mu\nu} \dot{z}^\mu \dot{z}^\nu} \delta^{(2)}(x - z(\theta)) d\theta \quad (29)$$