Joint evolution of a Lorentz-covariant scalar field and its point charge in one space dimension

Presented by Lawrence Frolov. Based on work with A. Shadi Tahvildar-Zadeh and Samuel Leigh.

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> > September 16, 2024

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• On n + 1 dimensional space-time  $\mathbb{R}^{1,n} \cong \mathbb{R} \times \mathbb{R}^n$  with metric  $\eta = \text{diag}(1, -1, -1...)$  and coordinates  $x = (x^0, x^1, x^2, ...)$ .

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- A point charge is represented by a future directed time-like curve ("world-line")  $z(\theta) : \mathbb{R} \to \mathbb{R}^{1,n}$ ,  $\frac{dz^0}{d\theta} > |\frac{d\vec{z}}{d\theta}|$ .

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- Associated to the particle is a mass m > 0, scalar charge  $a \in \mathbb{R}$ , velocity  $u^{\mu} := \frac{dz^{\mu}}{d\tau}$ , and momentum  $p^{\mu}$ .
- Notation: Jump and Average:

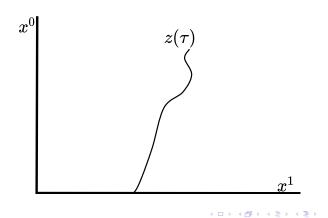
$$\overline{f}(z) = \lim_{\epsilon \to 0} \frac{f(z+\epsilon) + f(z-\epsilon)}{2}, \quad [f]_{x=z} = \lim_{\epsilon \to 0^+} \left(f(z+\epsilon) - f(z-\epsilon)\right)$$
(1)

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# Scalar Point Charge

The scalar point charge acts as a source for the scalar field via

$$\Box U := \partial_0^2 U - \Delta U = a \int \delta^{(n+1)}(x - z(\tau)) d\tau$$
(2)



# Motion of Scalar Field Animation

Given initial data  $(V_0, V_1)$  and a charge trajectory  $z(\tau)$ , the solution for U is well-known.

$$\begin{cases} \Box U = a \int \delta^{(n+1)} (x - z(\tau)) d\tau \\ U \big|_{x^0 = 0} = V_0 \\ \partial_0 U \big|_{x^0 = 0} = V_1 \end{cases}$$
(3)

Figure: Animation of scalar field sourced by a moving point charge in one space dimension.

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$$T_{p}^{\mu\nu} := \int u^{\mu} p^{\nu} \delta^{(n+1)}(x - z(\tau)) d\tau$$
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$$(1 + \frac{1}{\kappa^2} \Box) \Box U = J[z] \qquad \text{K-T.Z[2], Hoang et al.[3]}$$
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• In one space dimension, no modification is necessary!

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For initial data smooth away from z(0), this force can be explicitly computed as

$$\frac{dp^{\nu}}{d\tau} = -a\overline{\partial^{\nu}U}.$$
(12)

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$$\int_{\Omega} \partial_{\mu} T^{\mu\nu} dV = \int_{\partial \Omega} T^{\mu\nu} N_{\mu} dS$$
(13)

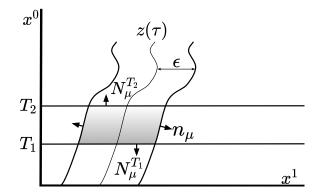
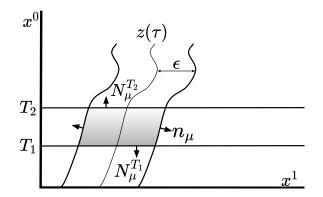
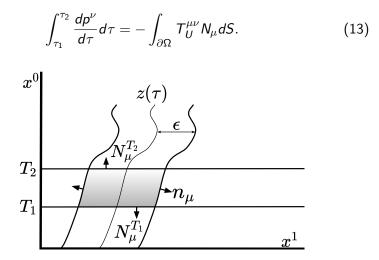


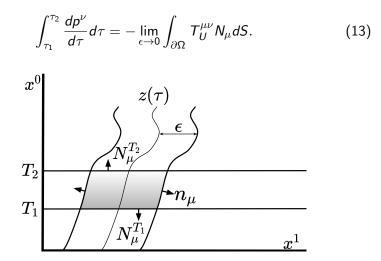
Figure: Region of integration  $\Omega$  and its normals

$$\int_{\partial\Omega} T_p^{\mu\nu} N_\mu dS = p^\nu(\tau_2) - p^\nu(\tau_1) = \int_{\tau_1}^{\tau_2} \frac{dp^\nu}{d\tau} d\tau.$$
(13)



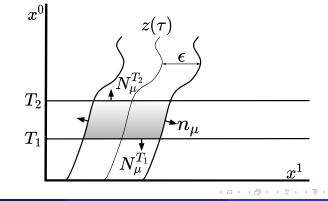
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#### Lemma

Suppose U satisfies equation (3) with  $(U, \partial_0 U)|_{x^0=0} \in C^{0,1}(\mathbb{R}) \times L^{\infty}(\mathbb{R})$ and singular source concentrated on a time-like world-line  $z(\tau)$ . Then U is of regularity  $C^{0,1}(\mathbb{R}^{1,1})$ , and  $\partial_{\mu}U, T_U^{\mu\nu} \in L^{\infty}(\mathbb{R}^{1,1})$ .



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$$\lim_{\epsilon \to 0} \int_{\partial \Omega} T_{U}^{\mu\nu} N_{\mu} dS = \int_{\tau_{1}}^{\tau_{2}} \left[ T_{U}^{\mu\nu} n_{\mu} (z^{0}, x^{1}) \right]_{x^{1} = z^{1}(\tau)} d\tau \qquad (14)$$

$$\begin{array}{c} x^{0} \\ T_{2} \\ T_{1} \\ \hline \end{array}$$

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$$\int_{\tau_1}^{\tau_2} \frac{dp^{\nu}}{d\tau} d\tau = -\int_{\tau_1}^{\tau_2} \left[ n_{\mu} T_U^{\mu\nu}(z^0, x^1) \right]_{x^1 = z^1(\tau)} d\tau.$$
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as desired.

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•  $n_{\mu}T^{\mu\nu}$  depends on both  $\partial^{0}U, \partial^{1}U$ , which seems undesireable since we'd like  $\frac{dp^{\nu}}{d\tau} \propto -a\partial^{\nu}U$ .

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- If we assume  $(U, \partial_0 U)|_{x^0=0}$  is smooth away from  $z^1(0)$ . We then have the jump formula

$$\left[\partial_{\mu}U\right]_{x^{1}=z^{1}(\tau)}=-an_{\mu}\tag{16}$$

The jump in the gradient is orthogonal to  $u^{\mu}$ !

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One can use this to compute

$$\frac{dp^0}{d\tau} = \frac{-1}{2u^1} \left[ (\partial_0 U)^2 \right]_{x^1 = z^1},\tag{18}$$

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The final force law  $\frac{dp^{
u}}{d au} = -a\overline{\partial^{
u}U}$  follows from

$$\left[\frac{f^2(x^1)}{2}\right]_{x^1=z^1} = [f]_{x^1=z^1} \overline{f}$$
(20)

#### Corollary

The force contribution of external scalar fields follows the standard law derived from the principle of least action.

$$F_{ext}^{\nu} = -a\partial^{\nu}U_{ext}.$$
 (21)

However, the singular "self-force" can now be determined, and the expression for it guarantees the conservation of the system's total energy-momentum.

$$F_{self}^{\nu} = -a\overline{\partial^{\nu}U}_{source} = \frac{-a^2u^{\nu}}{2}$$
(22)

#### Theorem

For any set of particle parameters with positive bare mass and non-zero real scalar charge, and for any "admissible" initial data, the joint initial value problem given by

$$\begin{cases} \eta^{\mu\nu}\partial_{\mu}\partial_{\nu}U(x) = a\int \delta^{(2)}(x-z(\tau))d\tau \\ U(0,x^{1}) = V_{0}(x^{1}) \\ \partial_{0}U(0,x^{1}) = V_{1}(x^{1}), \end{cases}$$
(23)

$$\begin{cases} \frac{dz^{\mu}}{d\tau} = u^{\mu} \\ \frac{dp^{\mu}}{d\tau} = -\left[n_{\mu}T_{U}^{\mu\nu}(z^{0}, x^{1})\right]_{x^{1}=z^{1}(\tau)}. \end{cases}$$
(24)

admits a unique, global-in-time solution.

• We studied the joint evolution equations for a scalar field and its point charge source in 1 space dimension

$$\Box U = a \int \delta^{(2)}(x - z(\tau)) d\tau.$$
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- We rigorously derived an equation of motion for the point charge from weak energy-momentum conservation

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- The force law closely resembles the one derived from PoLA, and returns a simple expression for the self-force.

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- We rigorously derived an equation of motion for the point charge from weak energy-momentum conservation
- The force law closely resembles the one derived from PoLA, and returns a simple expression for the self-force.
- This force law allowed us to prove well-posedness of the joint evolution problem for the field-particle singularity system.

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- Point mass acts as singularity in curvature

$$g_{\mu\nu} \in C^{0,1}(\mathcal{M}), \quad \left[\Gamma^{\kappa}_{\mu\nu}(z^0,x^1)\right]_{x^1=z^1} \neq 0$$
 (25)

$$\nabla_{\mu} T_{\rho}^{\mu\nu} \propto \partial_{\mu} T_{\rho}^{\mu\nu} + \Gamma T_{\rho} = 0?$$
(26)

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- In higher dimensions we need to study non-linear wave equations coupled to point sources.

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- In 1 space dimension, a non-linear theory of distributions will be necessary to study the gravitational joint evolution problem.
- In higher dimensions we need to study non-linear wave equations coupled to point sources.
- Non-linear modifications may regularize the sourced fields to permit a well-posed joint evolution

$$\partial_{\mu}\left(\frac{\partial^{\mu}U}{\sqrt{1+\partial_{\alpha}U\partial^{\alpha}U}}\right) = a\int \delta^{(n+1)}(x-z(\tau))d\tau.$$
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# White Lies

• True force law from Energy-Momentum conservation:

$$\int_{\tau_1}^{\tau_2} \frac{dp^{\nu}}{d\tau} d\tau = \lim_{\epsilon \to 0^+} \left( \int_{\tau_1}^{\tau_2} n_{\mu} T_U^{\mu\nu}(z^0, z^1 + \epsilon) - n_{\mu} T_U^{\mu\nu}(z^0, z^1 - \epsilon) d\tau \right)$$
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- The momentum of a scalar point charge is  $p^{\mu} := (m aU(z))u^{\mu} \neq mu^{\mu}$
- Action Principle:

$$S[U, z] = \int (\mathcal{L}_U + \mathcal{L}_i + \mathcal{L}_p) \sqrt{-\eta} dx^2$$
(27)

$$\mathcal{L}_{U} := \frac{1}{2} \eta^{\mu\nu} \partial_{\mu} U \partial_{\nu} U, \qquad \mathcal{L}_{i} := \frac{-a}{m} U(z) \mathcal{L}_{p}$$
(28)

$$\mathcal{L}_{p} := -\frac{m}{\sqrt{-\eta}} \int \sqrt{\eta_{\mu\nu} \dot{z}^{\mu} \dot{z}^{\nu}} \delta^{(2)}(x - z(\theta)) d\theta$$
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